# Requirements for a new EOS interface in the Einstein Toolkit

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**CCT: Center for Computation & Technology** 

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**Discussion Topic** 

- Topic: Requirements for a new EOS interface
- Background: Plan to restructure current EOS interface (EOSBase) for future needs
- Today: Collect / brainstorm requirements
- Will then discuss design and implementation by email



# Cactus EOS Interface

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XiRel phone call, 3 November 2009

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Requirements



Interfaces

- Original interface
- "New" interface





An EOS logically closes the system of equations.

Specifically, given some set of *independent* thermodynamic potentials (some subset of  $\rho$ ,  $\varepsilon$ , composition, entropy, temperature, ...) all remaining *dependent* quantities (pressure, speed of sound, entropy, ...) are given by the EOS.

The EOS must specify

- which independent variables it requires;
- which dependent variables it can compute.

# Original interface



In CactusEOS there exists the original EOS\_Base interface.

- Assumes that independent variables are *exactly* two of  $\rho, \varepsilon, P$ .
- Only potential dependent variables are  $\rho, \varepsilon, P, P_{\rho}, P_{\varepsilon}$ .
- Uses direct function pointers (pre-function aliasing).
- Calls not set up in advance no need.
- All calls are pointwise.
- Function prototypes fully specified e.g. press(i,j,k) = EOS\_Pressure(handle, rho(i,j,k), eps(i,j,k)).



In  ${\tt Whisky\_Dev}\ a\ new\ interface\ {\tt EOSBase\_General}\ exists\ but\ is\ little\ used.$ 

- (In)dependent variable numbers and names arbitrary.
- Uses function aliasing.
- Works on whole grid functions (∃ array interface, can be used for pointwise calls overly complex).
- Calls set up in advance and cached using tables.
- Function prototypes specified, but order of (in)dependent variables not e.g.

ierr = EOS\_SetGFs(cctkGH, EOS\_Con2PrimCall).
Here EOS\_Con2PrimCall is a table handle.

# Wishlist



EOS interface must be able to

- deal with arbitrary (in)dependent variables.
- work pointwise or with grid functions, and do it efficiently.
- be simple for simple cases.
- extend to table readers with fallback mechanisms.

### EOS interface should be able to

- deal with masks (excision, atmosphere) when using GFs.
- deal with multiple, possibly dynamically changing EOSs.
- provide a fallback mechanism for local failures.

EOS interface

### Roland Haas (Georgia Tech) Center for Relativistic Astrophysics

October 30, 2009

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EOS interface

October 30, 2009 1 / 4

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- Con2Prim: pointwise only; iterative
- Prim2Con: mostly for the whole grid, pointwise in AtmosphereReset
- Eigenproblem in Riemann solvers (Roe, Marquina): acts on whole grid
- (modified) PPM reconstruction: only pointwise interface used

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# Current EOS shortcomings

### old interface

- hard-coded dependend and independend variables
- EOSBase.h needs to be included in header of each (Fortran) function that uses EOS routines
- Cactus is not aware of the functions provided (not aliased)
- possible function call overhead at each point if whole grid function is processed (this should be small however)

### new (general) interface

- setup of calls is too complicated and error prone
- not publicly available (?)

- HydroBase: record EOS associated with primitives
- EOS interface should be simple to use
- keep number of temporary grid functions to a minimum
- maybe detect common errors at compile time (ie. misspelled variable names)
- fallback/alternative EOS mechanism (for Con2Prim)

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